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## GX 339-4

B. A. Harmon, G. J. Fishman, W. S. Paciesas, and M. Finger, Marshall Space Flight Center, NASA, report for the Burst and Transient Source Experiment (BATSE)/Compton Observatory Team: "The black-hole candidate GX 3394 has entered its hard (low) emission state, being detected by earth occultation beginning about mid-September. Its intensity has increased very gradually and is currently at ~ 0.24 Crab (Oct. 17-25) in the 20- to 300-keV band. The spectrum is very similar to the 1991 hard-state outburst (1A UC 5327), extending to at least 230 keV, with a power-law index of  $-2.1 \pm 0.06$ . A better fit is obtained, however, with a Comptonisation model, as the spectrum becomes steeper above 100 keV. The Sunyaev-Titarchuk model fit parameters for Oct. 17-25 are  $kT = 39 \pm 5$  keV and  $\tau = 2.5 \pm 0.5$ . Identification of GX 339-4 as the flaring source was obtained by fitting for the time of occultation edges, yielding a  $1\sigma$  error in location of  $\pm 0.35$  for  $\alpha$  and  $\pm 0.2$  for  $\delta$ . Monitoring of the source intensity and spectrum is continuing."

## JUPITER

G. Orton, P. Yanamandra-Fisher, and D. Griep, NASA Infrared Telescope Facility, report: "Eighteen-jtm images of Jupiter on Oct. 26.68-26.78 UT show that the South Equatorial Belt (SEB) has resumed a 150-mbar temperature equal to that of the NEB, recovering from prior cooler temperatures which lasted ~ 1 yr. Infrared indicators at 8.57 and 4.9  $\mu$ m show the SEB remaining unusually cloudy, as it has been since early 1992. If it behaves as in 1988-1990, the SEB warming signals the recovery of its cloud properties to a more typical state within 6-9 months."

## NOVASAGITTARII 1992 No. 2

Further photometry (cf. 1A UC 5633) by A. C. Gilmore, Mount John University Observatory: Oct. 17.42 UT,  $V = 12.62 \pm 0.08$ ,  $B - V = +0.21 \pm 0.08$ ,  $U - B = -0.26 \pm 0.08$ ,  $V - R = -1.12 \pm 0.09$

## NOVASAGITTARII 1992 No. 3

Further photometry (cf. 1A UC 5642) by Gilmore: Oct. 24.38 UT,  $V = 8.92 \pm 0.04$ ,  $B - V = +0.40 \pm 0.03$ ,  $U - B = -0.40 \pm 0.05$ ,  $V - R = -0.56 \pm 0.03$ ,  $V - I = +0.87 \pm 0.03$ .

1992 November 3

Daniel W. E. Green